

prob sheet 3.

Recall $u = t - r_*$ $du = dt - dr_*$
 $v = t + r_*$ $dv = dt + dr_*$

$$du dv = dt^2 - (dr_*)^2$$

$$= dt^2 - \left(1 - \frac{2M}{r}\right)^{-1} dr^2$$

$$\Rightarrow ds^2 = -\left(1 - \frac{2M}{r}\right) dt^2 + \left(1 - \frac{2M}{r}\right)^{-1} dr^2 + r^2 d\Omega$$

$$\boxed{ds^2 = -\left(1 - \frac{2M}{r}\right) du dv + r^2 d\Omega}$$

(note as an aside $g_{uv} = -\frac{1}{2}\left(1 - \frac{2M}{r}\right)$
 $\uparrow$ note the $1/2$!)

$$U = -e^{-u/4M} \Rightarrow dU = e^{-u/4M} \left(\frac{du}{4M}\right) = \frac{1}{4M} U du$$

$$V = e^{v/4M} \Rightarrow dV = \frac{1}{4M} V dv$$

$$\Rightarrow ds^2 = -\left(1 - \frac{2M}{r}\right) \left(\frac{dU}{\frac{U}{4M}}\right) \frac{4M}{V} dV + r^2 d\Omega$$

$$= -16M^2 \left(1 - \frac{2M}{r}\right) \frac{1}{UV} dU dV + r^2 d\Omega$$

Also have $UV = -e^{r_*/2M} = -e^{r/2M} (r/2M - 1)$
 $= -\frac{r}{2M} e^{r/2M} (1 - 2M/r)$

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$$\Rightarrow \boxed{ds^2 = -32M^3 \frac{e^{-r/2M}}{r} dU dV + r^2 d\Omega}$$

with $r = r(U, V)$.

in region I coord transn. $(t, r, \theta, \phi) \rightarrow (U, V, \theta, \phi)$

$$k = \partial_t = \frac{\partial U}{\partial t} \partial_U + \frac{\partial V}{\partial t} \partial_V + \cancel{\frac{\partial \theta}{\partial t} \partial_\theta} + \cancel{\frac{\partial \phi}{\partial t} \partial_\phi}$$

to work out $\frac{\partial U}{\partial t}$ & $\frac{\partial V}{\partial t}$ note that we are taking partial derivatives keeping r fixed. When r is fixed we have $U = \frac{c}{V}$ for constant c .

We also have $\frac{V}{U} = -e^{t/2M}$

$$\Rightarrow V^2 = -c e^{t/2M}$$

$$U^2 = -c e^{-t/2M}$$

$$\Rightarrow 2V \partial_t V = -c e^{t/2M} \frac{1}{2M} = +V^2/2M$$

$$\Rightarrow \partial_t V = V/4M$$

Also $2U \partial_t U = +c e^{-t/2M} \left(\frac{-1}{2M}\right) = -U^2/2M$

$$\Rightarrow \partial_t U = -U/4M$$

Hence

$$k = \partial_t = \frac{1}{4M} (V \partial_V - U \partial_U) //$$

In these coords $k^\mu = \frac{1}{4M} (-U, V, 0, 0)$

$$\|k\|^2 = k^\mu k^\nu g_{\mu\nu} = 2 k^U k^V g_{UV}$$

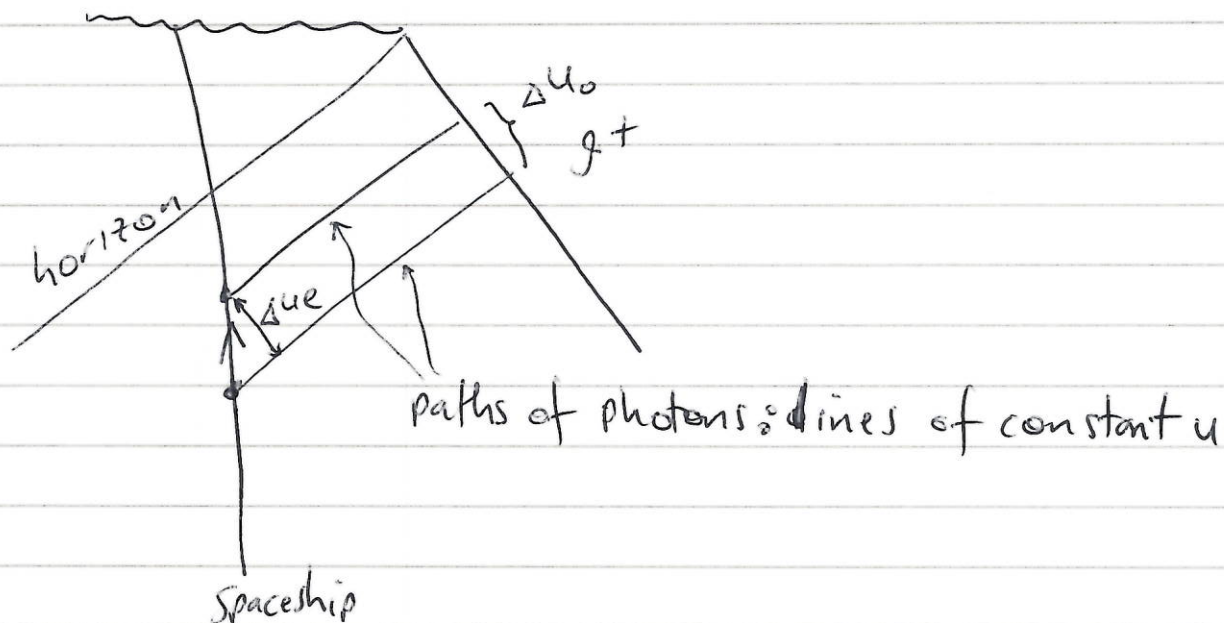
$$= 2 \frac{1}{(4M)^2} (-U)(V) \left(-\frac{16M^3}{r}\right) e^{-r/2M} = -\left(1 - \frac{2M}{r}\right)$$

In schwartschild coord's.

$$k^\mu = (1, 0, 0, 0)$$

$$\|k\|^2 = g_{\mu\nu} k^\mu k^\nu = g_{tt} = -(1 - 2M/r) \quad \checkmark$$

2.



Δu_e = difference in the u coordinate of 2 events where photons emitted

Δu_o = " " " received

But $\Delta u_e = \Delta u_o$ because photons travel along lines of constant u .

$$ds^2 = -\left(1 - \frac{2M}{r}\right) du^2 - 2 du dr + r^2 d\Omega$$

For observer at $r = \infty$, have constant r, θ, ϕ

$$\Rightarrow ds^2 = -d\tau_o^2 = -du^2$$

$$\Rightarrow \Delta\tau_o = \Delta u_o$$

The frequency of the photons at the emitter is ν_e
 " " " " " + " observer is ν_o

Redshift $\frac{v_e}{v_o} = \frac{(\Delta\tau_e)^{-1}}{(\Delta\tau_o)^{-1}} = \frac{\Delta\tau_o}{\Delta\tau_e} = \frac{\Delta u_e}{\Delta\tau_e} = \frac{du}{d\tau}\bigg|_e$

We want to calculate $\frac{du}{d\tau}\bigg|_e$ as spaceship

approaches $r=2M$ we will write $\delta = r-2M$

and neglect $o(\delta^2)$ terms. Note $r = 2M(1 + \frac{\delta}{2M})$

Now for spaceship

$$-1 = -\left(1 - \frac{2M}{r}\right) \left(\frac{du}{d\tau}\right)^2 - 2\frac{du}{d\tau} \frac{dr}{d\tau} \quad (*)$$

so we need $\frac{dr}{d\tau}$

Recall from lectures.

$$\left(\frac{dr}{d\tau}\right)^2 = (1-\epsilon^2) \left(\frac{R_{\max}}{r} - 1 \right) \quad R_{\max} = \frac{2M}{1-\epsilon^2}$$

$$= (1-\epsilon^2) \frac{1}{r} \left(\frac{2M}{(1-\epsilon^2)} - (1-\epsilon^2)r \right)$$

$$\approx \epsilon^2 \left(1 - \frac{\delta}{2M\epsilon^2} \right)$$

$$\Rightarrow \frac{dr}{d\tau} \approx \epsilon \left(1 - \frac{\delta}{4M\epsilon^2} \right)$$

Hence, from (*)

$$-\frac{\delta}{2M} \left(\frac{du}{d\tau}\right)^2 + 2\epsilon \left(1 - \frac{\delta}{4M\epsilon^2}\right) \frac{du}{d\tau} + 1 \approx 0$$

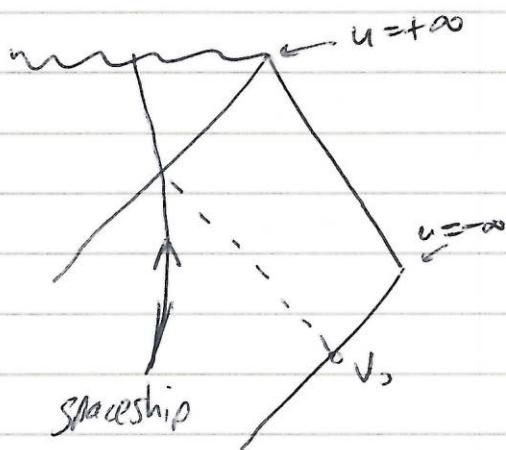
$$\Rightarrow \frac{du}{d\tau} = \frac{-2\epsilon \left(1 - \frac{\delta}{4M\epsilon^2}\right) \pm \left[4\epsilon^2 \left(1 - \frac{\delta}{4M\epsilon^2}\right)^2 + 4\frac{\delta}{2M} \right]^{1/2}}{-2\left(\frac{\delta}{2M}\right)}$$

expand square root & neglect $o(\delta^2)$:

$$\Rightarrow \frac{du}{dt} \approx \frac{-2\epsilon(1 - \delta/4M\epsilon^2) \pm (4\epsilon^2)^{1/2}}{-2(\delta/2M)}$$

$\frac{du}{dt} > 0$ so take lower sign

$$\Rightarrow \left. \frac{du}{dt} \right|_e \approx \frac{4M\epsilon}{\delta}$$



The emitter crosses horizon $r=2M$ at some value of v , v_0 say.

As $r \rightarrow 2M$

$$\Rightarrow v \rightarrow v_0$$

$$\Rightarrow t \rightarrow v_0 - r_*$$

$$\Rightarrow u \rightarrow v_0 - 2r_*$$

$$\Rightarrow u \rightarrow v_0 - 2\left(r + 2M \ln\left(\frac{r-2M}{2M}\right)\right)$$

and $u \rightarrow +\infty$ as $\delta \rightarrow 0$ since $\ln(-2M) = \ln \delta$

$$\Rightarrow u \approx v_0 - 2(2M) - 4M \ln(\delta/2M)$$

$$\approx v_0 - 4M + 4M \ln 2M - 4M \ln \delta$$

$$\text{or } u \approx \underbrace{\text{constant}}_{v_0} - 4M \ln \delta$$

$$\text{or } \delta^{-1} \approx e^{(u-v_0)/4M}$$

Thus $\left. \frac{du}{dt} \right|_e \approx 4M\epsilon e^{(u-v_0)/4M} = \text{redshift}$

Recall that u = proper time at infinity. Hence, redshift increases exponentially in proper time at ∞ and signals disappear in time of order $4M$. //

$$3. \quad x^0 = \cos \theta$$

$$x^\pm = \sin \theta \, \mu^\pm$$

$$(x^0)^2 + (x^\pm)^2 = \cos^2 \theta + \sin^2 \theta \, \underbrace{\mu^\pm \mu^\pm}_1 = \cos^2 \theta + \sin^2 \theta = 1$$

So parametrises S^N

choose

$0 \leq \theta \leq \pi$ so that we cover $-1 \leq x^0 \leq 1$

$$ds^2 = dx^0 dx^0 + dx^\pm dx^\pm$$

$$dx^0 = -\sin \theta \, d\theta$$

$$dx^\pm = \cos \theta \, d\theta \, \mu^\pm + \sin \theta \, d\mu^\pm$$

$$ds^2 = \sin^2 \theta \, d\theta^2 + [\cos^2 \theta \, d\theta^2 \, \mu^\pm \mu^\pm + \sin^2 \theta \, d\mu^\pm d\mu^\pm + 2\cos \theta \sin \theta \, d\theta \, \mu^\pm d\mu^\pm]$$

$$= \sin^2 \theta \, d\theta^2 + \cos^2 \theta \, d\theta^2 + \sin^2 \theta \, (d\Omega_{N-1})$$

$$+ 2\cos \theta \sin \theta \, d\theta \, \mu^\pm d\mu^\pm$$

Note that $\mu^\pm \mu^\pm = 1$

$$\Rightarrow 2\mu^\pm d\mu^\pm = 0$$

Hence

$$ds^2 = d\theta^2 + \sin^2 \theta \, d\Omega_{N-1}$$

4. As in lectures. Conformal compactification
 $(M, \hat{g})$ has

$$d\hat{S}^2 = -4d\tilde{u}d\tilde{v} + \sin^2(\tilde{v} - \tilde{u}) d\Omega$$

$$-\pi/2 \leq \tilde{u} \leq \tilde{v} \leq \pi/2$$

$$u = t - r \quad u = \tan \tilde{u}$$

$$v = t + r \quad v = \tan \tilde{v}$$

(i) eg r fixed & $t \rightarrow \infty \Leftrightarrow u, v \rightarrow \infty \Leftrightarrow \tilde{u} = \pi/2$
 $\tilde{v} = \pi/2$

Note that ~~the~~ the radius of the S^2 is
 $\sin(\tilde{v} - \tilde{u}) = 0$ Hence is a point

(-i) eg r fixed & $t \rightarrow -\infty \Leftrightarrow u, v \rightarrow -\infty \Leftrightarrow \tilde{u} = -\pi/2$
 $\tilde{v} = -\pi/2$

$\sin(\tilde{v} - \tilde{u}) = 0$ so this is also a point

(i^o) eg t finite & $r \rightarrow \infty \Leftrightarrow u \rightarrow -\infty \Leftrightarrow \tilde{u} = -\pi/2$
 $v \rightarrow +\infty \quad \tilde{v} = +\pi/2$

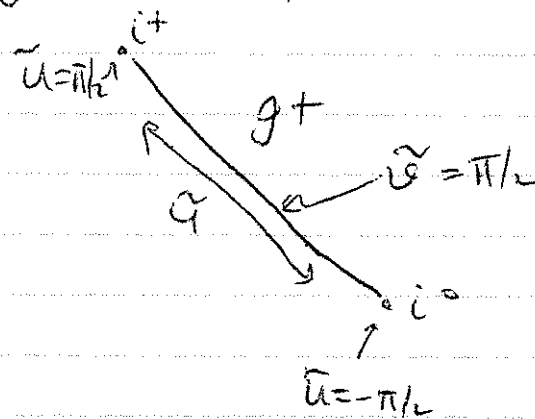
$\sin(\tilde{v} - \tilde{u}) = \sin(\pi) = 0$ and so again
 is a point.

(g^o) eg v finite & $u \rightarrow -\infty \Leftrightarrow \tilde{u} = -\pi/2$
 $|\tilde{v}| < \pi/2$

Notice $\sin(\tilde{v} - \tilde{u}) \neq 0$ Hence, I^+ is then
 product of the open interval $|\tilde{v}| < \pi/2$ with S^2
 or $\mathbb{R} \times S^2$

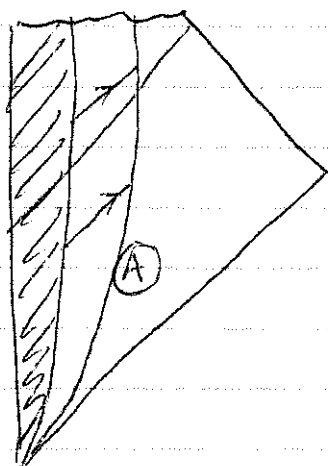
(g⁺) similar: u finite & $v \rightarrow +\infty : |\tilde{u}| < \pi/2$
 $|\tilde{v}| = \pi/2$

Notice, eg for $g+$ that as $\tilde{v} \rightarrow \pm\pi/L$ the

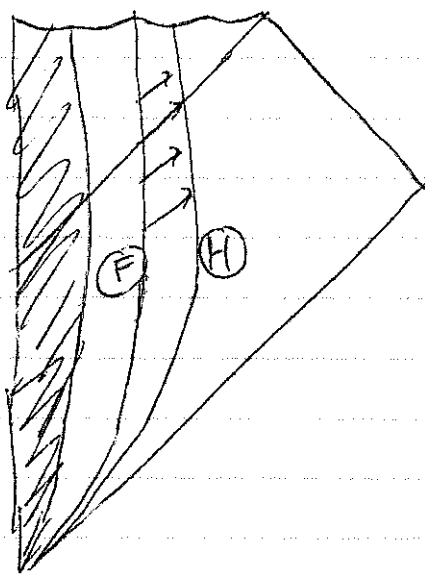


radius of the $S^2 \rightarrow 0$ and we reach the ~~AT~~ points i^+ & i^- , respectively.

5.



the arrowed lines represent radial geodesics from the star to the observer on worldline (A). yes, they still see the star.



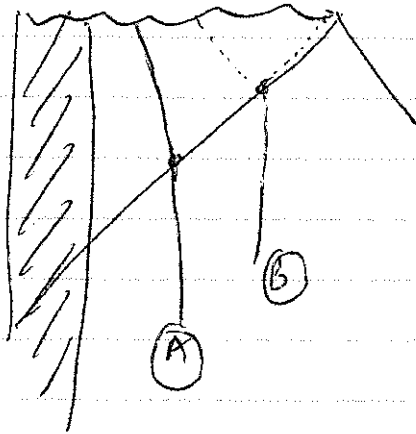
(H) represents worldline of head

(F) represents worldline of feet.

the arrowed lines are radial null geodesics from feet to the head. yes they see them.

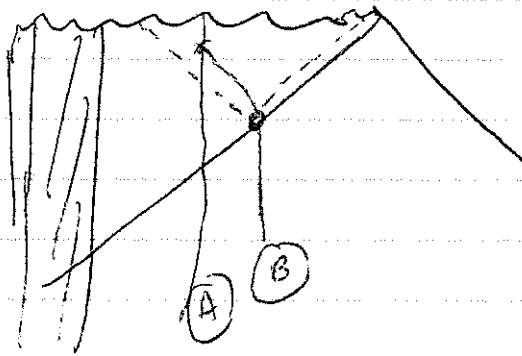
It depends.

~~Various possibilities~~

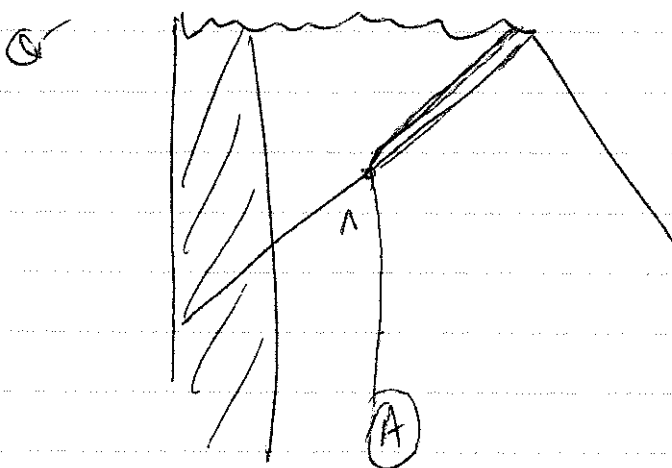


no matter what (B) does,
they won't meet (A)

On the other hand, they can meet



~~(Now, if (A) is accelerating)~~ accelerated



this worldline of
(A) will maximize
the possibility of
meeting.

$$\begin{aligned}
 \frac{1}{\sqrt{-g}} \partial_\mu \sqrt{-g} &= \frac{1}{\sqrt{-g}} \frac{1}{2\sqrt{-g}} \partial_\mu (-g) \\
 &= \frac{1}{2(-g)} (-g) g^{\rho\sigma} \partial_\mu g_{\rho\sigma} \\
 &= \frac{1}{2} g^{\rho\sigma} \partial_\mu g_{\rho\sigma}
 \end{aligned}$$

On the other hand we have

$$\begin{aligned}
 \Pi^\rho_{\mu} &= \frac{1}{2} g^{\rho\sigma} (g_{\sigma\rho,\mu} + \underbrace{g_{\sigma\mu,\rho} - g_{\rho\mu,\sigma}}_{\text{antisymmetric on } \rho \text{ and } \sigma}) \\
 &= \frac{1}{2} g^{\rho\sigma} \partial_\mu g_{\rho\sigma}
 \end{aligned}$$

Hence

$$\Pi^\rho_{\mu} = \frac{1}{\sqrt{-g}} \partial_\mu \sqrt{-g} //$$

$$\begin{aligned}
 \text{i) } \nabla_\mu V^\mu &= \partial_\mu V^\mu + \Pi^\mu_{\mu\sigma} V^\sigma \\
 &= \partial_\mu V^\mu + \left(\frac{1}{\sqrt{-g}} \partial_\mu \sqrt{-g} \right) V^\mu \\
 &= \frac{1}{\sqrt{-g}} \partial_\mu [\sqrt{-g} V^\mu] //
 \end{aligned}$$

$$\begin{aligned}
 \text{ii) } \nabla^2 \phi &\equiv g^{\mu\nu} \nabla_\mu \nabla_\nu \phi \\
 &= \nabla_\mu (g^{\mu\nu} \nabla_\nu \phi) \\
 &= \nabla_\mu (\nabla^\mu \phi) \\
 &= \frac{1}{\sqrt{-g}} \partial_\mu [\sqrt{-g} \nabla^\mu \phi] \\
 &= \frac{1}{\sqrt{-g}} \partial_\mu [\sqrt{-g} g^{\mu\nu} \partial_\nu \phi] //
 \end{aligned}$$

$$\begin{aligned}
 \text{iii) } \partial_\mu F^{\mu\nu} &= \partial_\mu F^{\mu\nu} + \underbrace{\Gamma_{\mu\sigma}^\mu}_{\text{symmetric}} F^{\sigma\nu} + \underbrace{\Gamma_{\mu\sigma}^\nu}_{\text{antisymmetric}} F^{\mu\sigma} \\
 &= \partial_\mu F^{\mu\nu} + \Gamma_{\mu\sigma}^\mu F^{\sigma\nu} \\
 &= \partial_\sigma F^{\sigma\nu} + \frac{1}{\sqrt{-g}} \partial_\sigma \sqrt{-g} F^{\sigma\nu} \\
 &= \frac{1}{\sqrt{-g}} \partial_\sigma [\sqrt{-g} F^{\sigma\nu}]
 \end{aligned}$$

Note it was crucial that $F_{\mu\nu} = -F_{\nu\mu}$ to get this cancellation & hence result.

$$\begin{aligned}
 \text{iv) } A_t &= Q/r \quad \text{and} \quad A_\phi = -P \cos \theta \\
 \Rightarrow F_{tr} &= \partial_t A_r - \partial_r A_t = -Q/r^2 \\
 F_{\theta\phi} &= \partial_\theta A_\phi - \partial_\phi A_\theta = P \sin \theta
 \end{aligned}$$

$$\begin{aligned}
 \text{Metric } \begin{cases} g_{tt} = -f & f = 1 - \frac{2M}{r} \\ g_{rr} = f^{-1} \\ g_{\theta\theta} = r^2 \\ g_{\phi\phi} = r^2 \sin^2 \theta \end{cases} \quad \begin{aligned} \det g &= -r^4 \sin^2 \theta \\ \Rightarrow \sqrt{-g} &= r^2 \sin \theta \end{aligned}
 \end{aligned}$$

$$\cancel{F^{\mu\nu}} F^{\mu\nu} = g^{\mu\sigma} g^{\nu\rho} F_{\sigma\rho}$$

$$\begin{aligned}
 F^{tr} &= g^{tt} g^{rr} F_{tr} = -F_{tr} = -Q/r^2 \\
 F^{\theta\phi} &= g^{\theta\theta} g^{\phi\phi} F_{\theta\phi} = \frac{P \sin \theta}{r^4 \sin^2 \theta} = \frac{P}{r^4 \sin \theta}
 \end{aligned}$$

$$\sqrt{-g} F^{tr} = r^2 \sin \theta \left(-\frac{Q}{r^2} \right) = -Q \sin \theta$$

$$\sqrt{-g} F^{\theta\phi} = r^2 \sin \theta \frac{P}{r^4 \sin \theta} = \frac{P}{r^2}$$

Max well's eq'ns $\nabla_\mu F^{\mu\nu} = 0$

$$\Leftrightarrow \partial_\mu (\sqrt{-g} F^{\mu\nu}) = 0$$

$$\begin{aligned} \underline{v=t} \quad \partial_\mu (\sqrt{-g} F^{\mu t}) &= \partial_r (\sqrt{-g} F^{rt}) \\ &= \partial_r (+Q \sin\theta) = 0 \end{aligned}$$

$$\underline{v=r} \quad \partial_\mu (\sqrt{-g} F^{\mu r}) = \partial_t (\sqrt{-g} F^{tr}) = 0$$

$$\begin{aligned} \underline{v=\theta} \quad \partial_\mu (\sqrt{-g} F^{\mu\theta}) &= \partial_\phi (\sqrt{-g} F^{\phi\theta}) \\ &= \partial_\phi (-p/r^2) = 0 \end{aligned}$$

$$\begin{aligned} \underline{v=\phi} \quad \partial_\mu (\sqrt{-g} F^{\mu\phi}) &= \partial_\theta (\sqrt{-g} F^{\theta\phi}) \\ &= \partial_\theta (p/r^2) = 0 \quad // \end{aligned}$$

Max well's eq'ns $\nabla_\mu F^{\mu\nu} = 0$

$$\Leftrightarrow \partial_\mu (\sqrt{-g} F^{\mu\nu}) = 0$$

$$\begin{aligned} \underline{v=t} \quad \partial_\mu (\sqrt{-g} F^{\mu t}) &= \partial_r (\sqrt{-g} F^{rt}) \\ &= \partial_r (+Q \sin\theta) = 0 \end{aligned}$$

$$\underline{v=r} \quad \partial_\mu (\sqrt{-g} F^{\mu r}) = \partial_t (\sqrt{-g} F^{tr}) = 0$$

$$\begin{aligned} \underline{v=\theta} \quad \partial_\mu (\sqrt{-g} F^{\mu\theta}) &= \partial_\phi (\sqrt{-g} F^{\phi\theta}) \\ &= \partial_\phi (-p/r^2) = 0 \end{aligned}$$

$$\begin{aligned} \underline{v=\phi} \quad \partial_\mu (\sqrt{-g} F^{\mu\phi}) &= \partial_\theta (\sqrt{-g} F^{\theta\phi}) \\ &= \partial_\theta (p/r^2) = 0 \quad // \end{aligned}$$

(v)

$$\begin{aligned}\nabla_\mu S^\mu_\nu &= \partial_\mu S^\mu_\nu + \Gamma^\mu_{\lambda\mu} S^\lambda_\nu - \Gamma^\lambda_{\mu\nu} S^\mu_\lambda \\&= \partial_\mu S^\mu_\nu + \frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g}) S^\mu_\nu - \Gamma^\lambda_{\mu\nu} S^\mu_\lambda \\&= \frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} S^\mu_\nu) - \Gamma^\lambda_{\mu\nu} S^\mu_\lambda\end{aligned}$$

Now

$$\begin{aligned}\Gamma^\lambda_{\mu\nu} S^\mu_\lambda &= \frac{1}{2} g^{\lambda\lambda} (g_{\lambda\mu,\nu} + g_{\lambda\nu,\mu} - g_{\mu\nu,\lambda}) S^\mu_\lambda \\&= \frac{1}{2} \underbrace{(g_{\lambda\mu,\nu} + g_{\lambda\nu,\mu} - g_{\mu\nu,\lambda})}_{\text{antisymmetric on } \mu \text{ and } \lambda} \underbrace{S^\mu_\lambda}_{\text{symmetric on } \mu \text{ and } \lambda} \\&= \frac{1}{2} g_{\lambda\mu,\nu} S^{\mu\lambda}\end{aligned}$$

$$\Rightarrow \nabla_\mu S^\mu_\nu = \frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} S^\mu_\nu) - \frac{1}{2} \partial_\nu (g_{\lambda\lambda}) S^{\lambda\mu} //$$

7.

Schwarzschild

$$ds^2 = -\left(1 - \frac{2M}{r}\right) dt^2 + 2dt dr + r^2 d\Omega$$

$$\tilde{t} = t - r$$

~~$$dt = d\tilde{t} + dr$$~~

$$d\tilde{t} = dt - dr$$

$$dt = (d\tilde{t} + dr)$$

$$ds^2 = -\left(1 - \frac{2M}{r}\right) (d\tilde{t}^2 + dr^2 + 2d\tilde{t} dr) + 2d\tilde{t} dr + 2dr^2 + r^2 d\Omega$$

~~$$ds^2 = -\left(1 - \frac{2M}{r}\right) (d\tilde{t}^2 + dr^2 + 2d\tilde{t} dr) + 2d\tilde{t} dr + 2dr^2 + r^2 d\Omega$$~~

$$= -\left(1 - \frac{2M}{r}\right) (d\tilde{t}^2 + dr^2 + 2d\tilde{t} dr) + 2d\tilde{t} dr + 2dr^2 + r^2 d\Omega$$

$$ds^2 = -d\tilde{t}^2 + dr^2 + r^2 d\Omega + \frac{2M}{r} (d\tilde{t} + dr)^2$$

We now introduce cartesian coords $\begin{cases} x + iy = r e^{i\phi} \sin\theta \\ z = r \cos\theta \end{cases}$

~~$$dx + i dy = (dr + i r d\phi) e^{i\phi}$$~~

~~$$dx^2 + dy^2 = dr^2 +$$~~

$$dx + i dy = (dr \sin\theta + r \cos\theta d\theta + r \sin\theta i d\phi) e^{i\phi}$$

$$dz = dr \cos\theta - r \sin\theta d\theta$$

$$\begin{aligned}
 dx^2 + dy^2 &= dr^2 \sin^2 \theta + r^2 \cos^2 \theta d\theta^2 + 2r \cos \theta \sin \theta dr d\theta \\
 &\quad + r^2 \sin^2 \theta d\phi^2 \\
 &\quad + dr^2 \cos^2 \theta + r^2 \sin^2 \theta d\theta^2 - 2r \sin \theta \cos \theta dr d\theta \\
 &= dr^2 + r^2 (d\theta^2 + \sin^2 \theta d\phi^2) \\
 &= dr^2 + r^2 d\Omega
 \end{aligned}$$

$$ds^2 = -d\tilde{t}^2 + dx^2 + dy^2 + dz^2 + \frac{2m}{r} (d\tilde{t} + dr)^2$$

$$r = r(x, y, z) \quad \text{via} \quad r^2 = x^2 + y^2 + z^2$$

$$\Rightarrow dr = \frac{x}{r} dx + \frac{y}{r} dy + \frac{z}{r} dz$$

~~$$ds^2 = -d\tilde{t}^2 + dx^2 + dy^2 + dz^2 + \frac{2m}{r} (d\tilde{t} + dr)^2$$~~

In these coord's

$$g_{\mu\nu} = \eta_{\mu\nu} + f K_\mu K_\nu$$

$$\text{where } K_\mu = \left(1, \frac{x}{r}, \frac{y}{r}, \frac{z}{r}\right) \quad \& \quad f = \frac{2m}{r}$$

First check conditions with respect to $\eta_{\mu\nu}$:

$$1) \text{ Notice } K^2 = -1 + \frac{x^2}{r^2} + \frac{y^2}{r^2} + \frac{z^2}{r^2} = 0 \quad \text{Null}$$

$$\begin{aligned}
 2) \quad K^\mu \nabla_\mu K_\nu &= K^\mu \partial_\mu K_\nu \\
 &= \cancel{\partial_t K_\nu} + \frac{x}{r} \partial_x K_\nu + \frac{y}{r} \partial_y K_\nu + \frac{z}{r} \partial_z K_\nu \\
 &= 0
 \end{aligned}$$

$$\underline{u=t}: K^{\mu} \partial_{\mu} K_t = 0$$

$$K_u = 1$$

$$\underline{u=x}: K_x = \frac{x}{r}$$

$$\begin{aligned} K^{\mu} \partial_{\mu} K_x &= \frac{x}{r} \partial_x \left(\frac{x}{r} \right) + \frac{y}{r} \partial_y \left(\frac{x}{r} \right) + \frac{z}{r} \partial_z \left(\frac{x}{r} \right) \\ &= \frac{x}{r} \left(\frac{1}{r} - \frac{x}{r^2} \frac{x}{r} \right) + \frac{yx}{r} \left(\frac{-1}{r^2} \right) \frac{y}{r} + \frac{zx}{r} \left(\frac{-1}{r^2} \right) \frac{z}{r} \\ &= \frac{x}{r^2} - \frac{x^3}{r^4} - \frac{xy^2}{r^4} - \frac{xz^2}{r^4} \\ &= \frac{x}{r^2} - x \frac{r^2}{r^4} \\ &= 0 \end{aligned}$$

Simil only for $\mu = y$ & z by symmetry.

Finally, check conditions with respect to $g_{\mu\nu}$.

$$\text{Firstly } g_{\mu\nu} = \eta_{\mu\nu} - f K^{\mu} K^{\nu}$$

$$\text{where } K^{\mu} = \eta^{\mu\nu} K_{\nu}$$

$$\begin{aligned} \text{check } g^{\mu\nu} g_{\nu\sigma} &= (\eta^{\mu\nu} - f K^{\mu} K^{\nu}) (\eta_{\nu\sigma} + f K_{\nu} K_{\sigma}) \\ &= \delta^{\mu}_{\sigma} + f K^{\mu} K_{\sigma} - f K^{\mu} K_{\sigma} + 0 \\ &= \delta^{\mu}_{\sigma} \quad \checkmark \end{aligned}$$

$$\text{thus } K^{\mu} = \eta^{\mu\nu} K_{\nu} = g^{\mu\nu} K_{\nu} \quad \text{also.}$$

$$\begin{aligned} K^2 &= g^{\mu\nu} K_{\mu} K_{\nu} \\ &= (\eta^{\mu\nu} - f K^{\mu} K^{\nu}) (K_{\mu}) (K_{\nu}) \\ &= \eta^{\mu\nu} K_{\mu} K_{\nu} \\ &= 0 \quad \checkmark \end{aligned}$$

$$K^{\mu} \nabla_{\mu} K_{\nu} = K^{\mu} (\partial_{\mu} K_{\nu} - \Gamma_{\mu\nu}^{\sigma} K_{\sigma})$$

We have shown $K^{\mu} \partial_{\mu} K_{\nu} = 0$ already.
So, just have to show

$$K^{\mu} K_{\sigma} \Gamma_{\mu\nu}^{\sigma} = 0$$

$$\Leftrightarrow K^{\mu} K_{\sigma} \frac{1}{2} g^{\sigma\delta} (g_{\delta\mu,\nu} + g_{\delta\nu,\mu} - g_{\mu\nu,\delta}) = 0$$

$$\Leftrightarrow \underbrace{K^{\mu} K^{\delta}}_{\text{symmetric in } \mu \text{ \& } \delta} (g_{\delta\mu,\nu} + \underbrace{g_{\delta\nu,\mu} - g_{\mu\nu,\delta}}_{\text{antisymmetric in } \mu \text{ \& } \delta}) = 0$$

$$\Leftrightarrow K^{\mu} K^{\delta} \frac{\partial}{\partial x^{\mu}} (g_{\delta\mu}) = 0$$

$$\Leftrightarrow K^{\mu} K^{\delta} \frac{\partial}{\partial x^{\mu}} (\eta_{\delta\mu} + f K_{\delta} K_{\mu}) = 0$$

$$\Leftrightarrow K^{\mu} K^{\delta} \frac{\partial}{\partial x^{\mu}} (f K_{\delta} K_{\mu}) = 0$$

& this is true
since $K^2 = 0 //$.

(i)

$$\begin{aligned}
 (A \wedge B)_{\mu_1, \dots, \mu_p, \nu_1, \dots, \nu_q} &= \frac{(p+q)!}{p!q!} A[\mu_1, \dots, \mu_p, \nu_1, \dots, \nu_q] \\
 &= \frac{(p+q)!}{p!q!} \left\{ A_{\mu_1, \dots, \mu_p, \nu_1, \dots, \nu_q} + \dots \right. \\
 &= \frac{1}{p!q!} \left\{ B_{\nu_1, \dots, \nu_q, \mu_1, \dots, \mu_p} + \dots \right. \\
 &= \frac{(p+q)!}{p!q!} B[\nu_1, \dots, \nu_q, \mu_1, \dots, \mu_p] \\
 &= (B \wedge A)_{\nu_1, \dots, \nu_q, \mu_1, \dots, \mu_p} \\
 &= (B \wedge A)_{\mu_1, \dots, \mu_p, \nu_1, \dots, \nu_q} (-1)^{q \cdot p}
 \end{aligned}$$

all perm's

we move e.g. ν_q past p of the μ_i + then do that for the rest of the $\nu_2, \dots, \nu_q$

$$\Rightarrow A \wedge B = (-1)^{pq} B \wedge A //$$

(ii)

$$\begin{aligned}
 d(A)_{\nu_1, \dots, \nu_{p+2}} &= (p+2) \partial_{\nu_1} dA_{\nu_2, \dots, \nu_{p+2}} \\
 &= (p+2)(p+1) \partial_{\nu_1} \partial_{\nu_2} A_{\nu_3, \dots, \nu_{p+2}}
 \end{aligned}$$

= 0 since $\partial_{\nu_1} \partial_{\nu_2}$ is symmetric on interchange of ν_1, ν_2

Note in the first step we know dA is a $(p+1)$ -form and we take the d of that object.

(iii)

$$(A \wedge B)_{\mu_1, \dots, \mu_p, \nu_1, \dots, \nu_q} = \frac{(p+q)!}{p!q!} A[\mu_1, \dots, \mu_p] B_{\nu_1, \dots, \nu_q}$$

~~Wanted to show~~

$$d(A \wedge B)_{\mu_1, \dots, \mu_p, \nu_1, \dots, \nu_q} = (p+q+1) \partial_{\mu_p} (A \wedge B)_{\mu_1, \dots, \mu_p, \nu_1, \dots, \nu_q}$$

$$= (p+q+1) \partial_{\mu_p} \left(\frac{(p+q)!}{p!q!} A[\mu_1, \dots, \mu_p] B_{\nu_1, \dots, \nu_q} \right)$$

$$= (p+q+1) \frac{(p+q)!}{p!q!} \left(\partial_{\mu_p} A[\mu_1, \dots, \mu_p] B_{\nu_1, \dots, \nu_q} + A[\mu_1, \dots, \mu_p] \partial_{\mu_p} B_{\nu_1, \dots, \nu_q} \right)$$

$$= \frac{(p+q+1)!}{p!q!} \left(\frac{(p+1)!q!}{(p+q+1)!} \frac{1}{(p+1)} (dA \wedge B)_{\mu_1, \dots, \mu_p, \nu_1, \dots, \nu_q} + \frac{p!(q+1)!}{(p+q+1)!} \frac{1}{q+1} (A \wedge dB)_{\mu_1, \dots, \mu_p, \nu_1, \dots, \nu_q} \right)$$

Where we call $(dA)_{\mu_1, \dots, \mu_p} = (p+1) \partial_{\mu_p} A[\mu_1, \dots, \mu_p]$

$$(dB)_{\nu_1, \dots, \nu_q} = (q+1) \partial_{\nu_q} B_{\nu_1, \dots, \nu_q}$$

$$= (dA \wedge B)_{\mu_1, \dots, \mu_p, \nu_1, \dots, \nu_q} + (A \wedge dB)_{\mu_1, \dots, \mu_p, \nu_1, \dots, \nu_q}$$

$$(-1)^p (A \wedge dB)_{\mu_1, \dots, \mu_p, \nu_1, \dots, \nu_q}$$

$$\Rightarrow A \wedge B = dA \wedge B + (-1)^p A \wedge dB //$$

(iv)

$$(* * A)_{\mu_1 \dots \mu_p}$$

Now $(*A)$ is a $(n-p)$ form, so take $*$ of that:

$$* (*A)_{\mu_1 \dots \mu_p} = \frac{1}{(n-p)!} \epsilon_{\mu_1 \dots \mu_p}^{\mu_{p+1} \dots \mu_n} (*A)_{\mu_{p+1} \dots \mu_n}$$

$$= \frac{1}{(n-p)!} \epsilon_{\mu_1 \dots \mu_p}^{\mu_{p+1} \dots \mu_n} \frac{1}{p!} \epsilon_{\mu_{p+1} \dots \mu_n}^{\nu_1 \dots \nu_p} A_{\nu_1 \dots \nu_p}$$

$$= \frac{1}{(n-p)!} \frac{1}{p!} \epsilon_{\mu_1 \dots \mu_p}^{\mu_{p+1} \dots \mu_n} (-1)^{p \cdot (n-p)} \epsilon_{\mu_{p+1} \dots \mu_n}^{\nu_1 \dots \nu_p} A_{\nu_1 \dots \nu_p}$$

↑

"Moved p indices past $n-p$ "

$$= \frac{1}{(n-p)!} \frac{1}{p!} (-1)^{p \cdot (n-p)} \cdot (\pm) p! (n-p)! \delta_{\mu_1 \dots \mu_p}^{\nu_1 \dots \nu_p} A_{\nu_1 \dots \nu_p}$$

$$= \pm (-1)^{p(n-p)} A_{\mu_1 \dots \mu_p}$$

$$\Rightarrow * * A = \pm (-1)^{p(n-p)} A //$$

Again Work "from the left" i.e. take $\times$ of the $(n-p+1)$ -form $(d \times A)$:

$$(V) (\times d \times A)_{\mu_1 \dots \mu_{p-1}}$$

$$= \frac{1}{(n-p+1)!} \sum_{\mu_1 \dots \mu_{p-1}} \int_{V_1 \dots V_{n-p}} (d \times A)_{\rho_1 \dots \rho_{n-p}}$$

$$= \frac{1}{(n-p+1)!} \sum_{\mu_1 \dots \mu_{p-1}} \int_{V_1 \dots V_{n-p}} (n-p+1) \nabla_{[\rho} \times A_{\nu_1 \dots \nu_{n-p}]}$$

$$= \frac{1}{(n-p)!} \sum_{\mu_1 \dots \mu_{p-1}} \int_{V_1 \dots V_{n-p}} \nabla_{\rho} \times A_{\nu_1 \dots \nu_{n-p}}$$

$$= \frac{1}{(n-p)!} \sum_{\mu_1 \dots \mu_{p-1}} \int_{V_1 \dots V_{n-p}} \nabla_{\rho} \left(\frac{1}{p!} \sum_{\sigma_1 \dots \sigma_p} \epsilon_{\sigma_1 \dots \sigma_p} A_{\sigma_1 \dots \sigma_p} \right)$$

$$= \frac{1}{(n-p)!} \frac{1}{p!} \sum_{\mu_1 \dots \mu_{p-1}} \int_{V_1 \dots V_{n-p}} \sum_{\nu_1 \dots \nu_{n-p}} \epsilon_{\nu_1 \dots \nu_{n-p}} \nabla_{\rho} A_{\sigma_1 \dots \sigma_p}$$

$$= \frac{1}{(n-p)!} \frac{1}{p!} (-1)^{(n-p)p} \sum_{\mu_1 \dots \mu_{p-1}} \sum_{\nu_1 \dots \nu_{n-p}} \epsilon_{\nu_1 \dots \nu_{n-p}} \nabla^{\rho} A_{\sigma_1 \dots \sigma_p}$$

$$\pm A! (n-p)! \delta_{\mu_1 \dots \mu_{p-1}}^{\sigma_1 \dots \sigma_{p-1}, \rho}$$

$$= \pm (-1)^{(n-p)p} \nabla^{\rho} A_{\mu_1 \mu_2 \dots \mu_{p-1} \rho}$$

~~THE MAXIMUM VALUE OF THE FUNCTION~~

$$\begin{aligned}
 \text{9. (i)} \quad g_{\mu\nu} g^{\nu\sigma} &= \delta_{\mu}^{\sigma} \Rightarrow \delta g_{\mu\nu} g^{\nu\sigma} + g_{\mu\nu} \delta g^{\nu\sigma} = 0 \\
 &\Rightarrow \delta g_{\mu\nu} g^{\nu\sigma} = -g_{\mu\nu} \delta g^{\nu\sigma} \\
 &\Rightarrow \delta g_{\mu\rho} = -g_{\mu\nu} g_{\rho\sigma} \delta g^{\nu\sigma} //
 \end{aligned}$$

$$\text{(ii)} \quad \Gamma_{\nu\rho}^{\mu} = \frac{1}{2} g^{\mu\sigma} (g_{\sigma\nu,\rho} + g_{\sigma\rho,\nu} - g_{\nu\rho,\sigma})$$

$$\begin{aligned}
 \delta \Gamma_{\nu\rho}^{\mu} &= \frac{1}{2} \delta g^{\mu\sigma} (g_{\sigma\nu,\rho} + g_{\sigma\rho,\nu} - g_{\nu\rho,\sigma}) \\
 &\quad + \frac{1}{2} g^{\mu\sigma} (\delta g_{\sigma\nu,\rho} + \delta g_{\sigma\rho,\nu} - \delta g_{\nu\rho,\sigma})
 \end{aligned}$$

$$\frac{1}{2} g^{\mu\sigma} (\delta g_{\sigma\nu;\rho} + \delta g_{\sigma\rho;\nu} - \delta g_{\nu\rho;\sigma})$$

$$+ \frac{1}{2} g^{\mu\sigma} \left(\cancel{\Gamma_{\sigma\rho}^{\delta} \delta g_{\delta\nu}} + \cancel{\Gamma_{\nu\sigma}^{\delta} \delta g_{\delta\rho}} - \cancel{\Gamma_{\sigma\nu}^{\delta} \delta g_{\delta\rho}} \right.$$

$$\left. + \Gamma_{\rho\nu}^{\delta} \delta g_{\sigma\delta} + \Gamma_{\nu\rho}^{\delta} \delta g_{\sigma\delta} - \cancel{\Gamma_{\sigma\rho}^{\delta} \delta g_{\nu\delta}} \right)$$

$$= \frac{1}{2} g^{\mu\sigma} (\delta g_{\sigma\nu;\rho} + \delta g_{\sigma\rho;\nu} - \delta g_{\nu\rho;\sigma})$$

$$+ \frac{1}{2} \delta g^{\mu\sigma} (g_{\sigma\nu,\rho} + g_{\sigma\rho,\nu} - g_{\nu\rho,\sigma})$$

$$+ \frac{g^{\mu\sigma} \delta g_{\sigma\delta} \Gamma_{\rho\nu}^{\delta}}{2}$$

$$- \delta g^{\mu\sigma} g_{\sigma\delta} \frac{1}{2} g^{\delta\tau} (g_{\tau\rho,\nu} + g_{\tau\nu,\rho} - g_{\rho\nu,\tau})$$

$$\Rightarrow \delta P^\mu_{\nu\rho} = \frac{1}{2} g^{\mu\sigma} (\delta g_{\sigma\nu;\rho} + \delta g_{\sigma\rho;\nu} - \delta g_{\nu\rho;\sigma}) //$$

$$(iii) \quad \delta R = \delta (R_{\mu\nu} g^{\mu\nu}) = (\delta R_{\mu\nu}) g^{\mu\nu} + R_{\mu\nu} \delta g^{\mu\nu}$$

$$R_{\nu\sigma} = R^\mu_{\nu\mu\sigma}$$

$$= \partial_\mu P^\mu_{\nu\sigma} - \partial_\sigma P^\mu_{\nu\mu} + P^\mu_\mu \delta P^\delta_{\nu\sigma} - P^\mu_\sigma \delta P^\delta_{\nu\mu}$$

$$\Rightarrow \delta R_{\nu\sigma} = \partial_\mu (\delta P^\mu_{\nu\sigma}) - \partial_\sigma (\delta P^\mu_{\nu\mu})$$

$$\nabla_\mu (\delta P^\mu_{\nu\sigma}) - \nabla_\sigma (\delta P^\mu_{\nu\mu})$$

$$- \left(P^\mu_\mu \delta P^\delta_{\nu\sigma} - P^\mu_{\nu\nu} \delta P^\delta_{\sigma\sigma} - P^\delta_{\mu\sigma} \delta P^\mu_{\nu\delta} \right.$$

$$\left. - P^\mu_\sigma \delta P^\delta_{\nu\mu} + P^\delta_{\sigma\nu} \delta P^\mu_{\delta\mu} + P^\delta_{\sigma\mu} \delta P^\mu_{\nu\delta} \right)$$

$$+ \delta P^\mu_\mu P^\delta_{\nu\sigma} + P^\mu_\mu \delta P^\delta_{\nu\sigma} - \delta P^\mu_\sigma P^\delta_{\nu\mu} - P^\mu_\sigma \delta P^\delta_{\nu\mu}$$

$$\Rightarrow \delta R_{\nu\sigma} = \nabla_\mu (\delta P^\mu_{\nu\sigma}) - \nabla_\sigma (\delta P^\mu_{\nu\mu})$$

$$\Rightarrow g^{\nu\sigma} \delta R_{\nu\sigma} = g^{\nu\sigma} \nabla_\mu (\delta P_{\nu\sigma}^\mu) - g^{\nu\sigma} \nabla_\sigma (\delta P_{\nu\mu}^\mu) \\ = \nabla_\mu (g^{\nu\sigma} \delta P_{\nu\sigma}^\mu) - \nabla^\nu (\delta P_{\nu\mu}^\mu)$$

$$= \nabla_\mu \left[g^{\nu\sigma} \frac{1}{2} g^{\mu\delta} (2 \delta g_{\delta\nu;\sigma} - \delta g_{\nu\sigma;\delta}) \right] \\ - \nabla^\nu \left[\frac{1}{2} g^{\mu\delta} (\delta g_{\delta\nu;\mu} + \delta g_{\delta\mu;\nu} - \delta g_{\mu\delta;\nu}) \right]$$

$$= \nabla^\delta \nabla^\nu \delta g_{\delta\nu} - \frac{1}{2} \nabla^\delta \nabla_\delta \delta g_{\nu\sigma} g^{\nu\sigma} \\ - \frac{1}{2} \nabla^\nu \nabla_\nu \delta g_{\delta\mu} g^{\mu\delta}$$

Since
2 terms
antisym.
in μ & δ

$$= \nabla^\delta \left[\underbrace{(\nabla^\nu \delta g_{\delta\nu}) - (\nabla_\delta \delta g_{\nu\sigma}) g^{\nu\sigma}}_{\equiv V_\delta} \right] \\ = \nabla^\delta V_\delta$$

(iv)

$$\delta \sqrt{-g} = \frac{1}{2} \frac{1}{\sqrt{-g}} (-g) g^{\mu\nu} \delta g_{\mu\nu} = \frac{1}{2} \sqrt{-g} g^{\mu\nu} \delta g_{\mu\nu} \\ = -\frac{1}{2} \sqrt{-g} g_{\mu\nu} \delta g^{\mu\nu}$$

(v)

$$\delta(\sqrt{-g} R) = (\delta\sqrt{-g}) R + \sqrt{-g} \delta R$$

$$= -\frac{1}{2} \sqrt{-g} g_{\mu\nu} \delta g^{\mu\nu} R + \sqrt{-g} [R_{\mu\nu} \delta g^{\mu\nu} + \nabla^\delta V_\delta]$$

$$= \sqrt{-g} \left(R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} \right) \delta g^{\mu\nu} + \sqrt{-g} \nabla^\mu V_\mu$$

$$\delta(\sqrt{-g} F_\mu F^{\mu\nu}) = (\delta\sqrt{-g}) F_\mu F^{\mu\nu} + \sqrt{-g} \delta(F^{\mu\nu} F_{\mu\nu})$$

$$= -\frac{1}{2} \sqrt{-g} g_{\mu\nu} \delta g^{\mu\nu} F^2$$

$$+ \sqrt{-g} \delta \left[F_{\rho\sigma} g^{\mu\rho} g^{\sigma\nu} F_{\mu\nu} \right]$$

$$\sqrt{-g} 2 \delta g^{\mu\rho} (F_{\rho\sigma} g^{\sigma\nu} F_{\mu\nu})$$

$$\sqrt{-g} 2 \delta g^{\mu\nu} (F_{\nu\sigma} F_\mu{}^\sigma)$$

$$= \sqrt{-g} \left[2 F_{\mu\sigma} F_\nu{}^\sigma - \frac{1}{2} F^2 g_{\mu\nu} \right] \delta g^{\mu\nu}$$

$$\mathcal{L} = \sqrt{-g} [R - F_{\mu\nu} F^{\mu\nu}]$$

$$\delta\mathcal{L} = \sqrt{-g} \left[R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} - 2 \left(F_{\mu\sigma} F_\nu{}^\sigma - \frac{1}{4} F^2 g_{\mu\nu} \right) \right] \delta g^{\mu\nu} + \sqrt{-g} \nabla^\mu V_\mu //$$